

Relativistic effects on light matter interactions

We continue to describe light semiclassically but improve our description of matter by considering relativistic effects. We will discover $\hbar\omega$ as an essentially relativistic attribute of elementary particles which resembles an intrinsic magnetic moment in the non-relativistic limit.

This chapter is devoted to understanding of how this intrinsic magnetic moment interacts with external magnetic moments due to relative motion of the ion and the electrons. We will learn how these interactions modify energy levels described by n, l and the associated selection rules.

First we will briefly review relativistic QM of free particles.

Relativistic relation E , momentum p and rest mass m :

$$E^2 = m^2 c^4 + p^2 c^2 \Rightarrow -\hbar^2 \frac{\partial^2 \psi}{\partial t^2} = m^2 c^4 \psi - \hbar^2 c^2 \nabla^2 \psi$$

\uparrow \uparrow \rightarrow Klein Gordon \hat{E}
 $i\hbar \frac{\partial}{\partial t}$ $-i\hbar \nabla$

Note that the KG equation is satisfied by $e^{i(kx \pm Et/\hbar)}$

$$\text{or } \psi = A e^{i(kx \pm Et/\hbar)}$$

$$\text{KG } \hat{E} \Rightarrow E^2 = m^2 c^4 + c^2 \hbar^2 k^2 \Rightarrow E \rightarrow \pm \infty \text{ as } k \rightarrow \infty$$

Note $e^{i(kx + Et/\hbar)}$ not allowed by KG. Since TDSE and TISE would give energies differing by sign.

However KG \hat{E} must be correct. So we need a different description for relativistic particles which will not lead to the apparently catastrophic scenario of E of particles going to $-\infty \rightarrow$ Global sink!

Dirac's proposal: $\hat{H} = c \vec{\alpha} \cdot \vec{p} + \beta mc^2$ | Relativistic $E = \pm \sqrt{p^2 c^2 + m^2 c^4}$

Motivation was to avoid double derivative of time but locate the spatial coordinates in the same footing at (or at more precisely) to be consistent with Sp. Relativity. The hope was that with suitable choice of $\vec{\alpha}$ and β the \hat{H} will merge into KG \hat{E} yet the description of particles will still be consistent with $E = \pm \sqrt{p^2 c^2 + m^2 c^4}$

$\therefore (E - c\vec{\alpha} \cdot \vec{p} - \beta mc^2)\Psi = 0 \rightarrow$ Dirac wave eq.
 with $i\hbar \frac{\partial \Psi}{\partial t} = E\Psi = H\Psi$

Let us consider Ψ not as a scalar field in configuration space but as column matrices for each point in coordinate sp.

(This generally becomes apparent a bit later.)

$\therefore H_{\vec{\alpha}, \beta}$ are $N \times N$ matrices and owing to the Hermiticity of $H_{\vec{\alpha}, \beta}$, $\vec{\alpha} = \vec{\alpha}^\dagger$ and $\beta = \beta^\dagger$

To find $\vec{\alpha}$ and β Dirac equated:

$(E + \vec{\alpha} \cdot \vec{p} + \beta mc^2)(E - c\vec{\alpha} \cdot \vec{p} - \beta mc^2)\Psi = 0$

to $[E^2 - \vec{p}^2 c^2 - m^2 c^4]\Psi = 0$ the KH eqⁿ

$$\left[E^2 - c^2 \left[\sum_{k=1}^3 (\alpha^k)^2 \hat{p}_k^2 + \sum_{(k < l)} (\alpha^k \alpha^l + \alpha^l \alpha^k) \hat{p}_k \hat{p}_l \right] - m^2 c^4 \right] \Psi = 0$$

Comparing above with the KH eqⁿ:

$$\left[\begin{aligned} (\alpha^1)^2 &= (\alpha^2)^2 = (\alpha^3)^2 = \beta^2 = 1 && \text{--- (A)} \\ [\alpha^1, \alpha^2]_+ &= [\alpha^2, \alpha^3]_+ = [\alpha^3, \alpha^1]_+ = 0 && \text{--- (B)} \\ [\alpha^1, \beta]_+ &= [\alpha^2, \beta]_+ = [\alpha^3, \beta]_+ = 0 && \text{--- (B)} \end{aligned} \right]$$

Note that $\{\alpha^k\}$ and β are scalars, then the set of (A) equation would have been enough but then they would not satisfy (B). So $\{\alpha^k\}$ and β must have more components in them and that leads to the matrix form. Now if we consider a 2×2 matrix for $\{\alpha^k\}$ and β , or more simply for $A^2 = B^2 = 1$ and $AB + BA = 0$ then we can show easily that it can start with σ_x

A possible (non unique) choice of $\alpha^1, \alpha^2, \alpha^3, \beta$ was readily available from the idea of quaternions.

$$\vec{\alpha} = \begin{pmatrix} 0 & \sigma^1 \\ \sigma^1 & 0 \end{pmatrix} \quad \vec{\alpha} = \hat{i}\sigma^1 + \hat{j}\sigma^2 + \hat{k}\sigma^3$$

\downarrow 2×2 \downarrow 2×2 \downarrow 2×2 \downarrow 2×2

$$\begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} \quad \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix} \quad \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}$$

and $\beta = \begin{pmatrix} I & 0 \\ 0 & -I \end{pmatrix}$; $I_2 = \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix}$

Analogously $\Psi = \begin{bmatrix} \psi_1 \\ \psi_2 \\ \psi_3 \\ \psi_4 \end{bmatrix}$, $\Psi^\dagger = (\psi_1^* \psi_2^* \psi_3^* \psi_4^*)$

$$\Psi(\vec{r}, t) = \Psi^\dagger \Psi = \sum_{i=1}^4 |\psi_i(\vec{r}, t)|^2$$

$$\Rightarrow \frac{\partial \Psi}{\partial t} + \nabla \cdot (\Psi^\dagger \vec{\alpha} \Psi) \quad (\text{show HW})$$

Charged particle in EM field:

$$[(E - q\phi) - c \vec{\alpha} \cdot (\vec{p} - q\vec{A}) - \beta mc^2] \Psi = 0$$

For time independent A and ϕ , as is the case for ^{quasi} atomic environment, due to a slowly moving ion, we can write

$$\Psi = \begin{bmatrix} \psi(\vec{r}) \\ \eta(\vec{r}) \end{bmatrix} e^{-iEt/\hbar} \quad \psi(\vec{r}) = \begin{bmatrix} \psi_1 \\ \psi_2 \end{bmatrix} \quad \eta(\vec{r}) = \begin{bmatrix} \psi_3 \\ \psi_4 \end{bmatrix}$$

$$E \psi(\vec{r}) = c(-i\hbar \nabla - q\vec{A}) \cdot \vec{\sigma} \eta(\vec{r}) + (q\phi + mc^2) \psi(\vec{r})$$

$$E \eta(\vec{r}) = c(-i\hbar \nabla - q\vec{A}) \cdot \vec{\sigma} \psi(\vec{r}) + (q\phi - mc^2) \eta(\vec{r})$$

Let $E = E' + mc^2$ and $|E'|$ as well as $|q\phi| \ll mc^2$

$$E' \psi(\vec{r}) = c(-i\hbar \nabla - q\vec{A}) \cdot \vec{\sigma} \eta(\vec{r}) + q\phi \psi(\vec{r}) \quad \text{--- (A)}$$

$$(E' + 2mc^2) \eta(\vec{r}) = c(-i\hbar \nabla - q\vec{A}) \cdot \vec{\sigma} \psi(\vec{r}) + q\phi \eta(\vec{r}) \quad \text{--- (B)}$$

considering only the dominant terms,
 Note (B1) $\hat{\Lambda} \rightarrow \eta(r) = \frac{1}{2mc} (-i\hbar\nabla - q\vec{A}) \cdot \vec{\sigma} \psi(r)$ (B1)

since the term with c or c^2 are far too greater than the other term.

$$\Rightarrow \frac{\eta(r)}{\psi(r)} \approx \frac{\hbar}{mc} \approx \frac{v}{c} \Rightarrow |\eta(r)| \ll |\psi(r)|$$

↓ weak ↓ strong

Substituting $\eta(r)$ from (B1) in (A1)

$$E'\psi = \frac{1}{2m} [(-i\hbar\nabla - q\vec{A}) \cdot \vec{\sigma}]^2 \psi + q\phi\psi$$

Use $(\vec{\sigma} \cdot \vec{D})(\vec{\sigma} \cdot \vec{B}) = \vec{D} \cdot \vec{B} + i\vec{\sigma} \cdot (\vec{D} \times \vec{B})$

but $\vec{D} = (-i\nabla - q\vec{A})$; $\vec{B} = (-\nabla - q\vec{A})\psi$

$$\therefore [(-i\hbar\nabla - q\vec{A}) \cdot \vec{\sigma}]^2 \psi = (-i\hbar\nabla - q\vec{A}) \cdot (-i\hbar\nabla - q\vec{A})\psi + i\vec{\sigma} \cdot [(-i\hbar\nabla - q\vec{A}) \times (-i\hbar\nabla - q\vec{A})\psi]$$

\downarrow
 $= i\hbar q (\nabla \times \vec{A}) \psi$

$$E'\psi = \frac{1}{2m} (-i\hbar\nabla - q\vec{A}) \cdot (-i\hbar\nabla - q\vec{A})\psi - \frac{\hbar q}{2m} \vec{\sigma} \cdot (\nabla \times \vec{A})\psi - e\phi\psi$$

$q = -e$

$$E'\psi = \left[\frac{1}{2m} (-i\hbar\nabla + e\vec{A}) \cdot (-i\hbar\nabla + e\vec{A}) + \frac{e\hbar}{2m} (\vec{\sigma} \cdot \vec{B}) - e\phi \right] \psi$$

(A2)

↳ Pauli Eqⁿ

Intrinsic magnetic moment: $\frac{e\hbar}{2m} \vec{\sigma} = \mu_B \vec{\sigma}$ conventionally: $(-\mu_B \vec{\sigma})$ → SPIN

The above equation is actually the Heuristically argued Pauli theory to explain Stern Gerlach type experiments, which showed the existence of SPIN. Non relativistic limit of Dirac theory is then the Pauli theory, which is two component. Recall that we are only talking about the 'strong' component.