

return to the Dirac theory to see the effect of spin at the relativistic level which we call "higher order" correction.

We will consider electron in Coulomb field of a nucleus. $\therefore A = 0$; $q\phi = -e\phi = V(r) = \frac{-Ze}{(4\pi\epsilon_0)r}$

Recall,

$$(B) \rightarrow (E' + 2mc^2)\underline{\eta}(r) = c(-i\hbar\nabla - \gamma A) \cdot \underline{\underline{\sigma}} \underline{\Psi}(r) + 2\phi \underline{\eta}(r)$$

$$(E' + 2mc^2)\underline{\eta}(r) = c(-i\hbar\nabla \cdot \underline{\underline{\sigma}}) \underline{\Psi} + V(r) \underline{\eta}(r)$$

$$\Rightarrow \underline{\eta}(r) = (E' + 2mc^2 - V(r))^{-1} c(-i\hbar \underline{\underline{\sigma}} \cdot \nabla) \underline{\Psi}(r)$$

Substituting the $\underline{\eta}(r)$ above in (A) with γ matrices A and ϕ .

Recall (A); $E'\underline{\Psi} = c(-i\hbar\nabla - \gamma A) \cdot \underline{\underline{\sigma}} \underline{\eta} + 2\phi \underline{\Psi}$

$$\rightarrow E'\underline{\Psi} = c^2 \frac{(-i\hbar\nabla \cdot \underline{\underline{\sigma}}) \cdot (-i\hbar \underline{\underline{\sigma}} \cdot \nabla) \underline{\Psi}}{(E' + 2mc^2 - V(r))} + V(r) \underline{\Psi}$$

Now, $\frac{1}{E' + 2mc^2 - V(r)} = \frac{1}{2mc^2(1 + \frac{E' - V(r)}{2mc^2})} = \frac{1}{2mc^2} \left[1 - \frac{E' - V(r)}{2mc^2} \right]$

$$\therefore E'\underline{\Psi} = c^2 (-i\hbar \nabla \cdot \underline{\underline{\sigma}}) \frac{1}{2mc^2} \left[1 - \frac{E' - V(r)}{2mc^2} \right] (-i\hbar \underline{\underline{\sigma}} \cdot \nabla) \underline{\Psi} + V(r) \underline{\Psi}$$

$$= \frac{1}{2m} \left[1 - \frac{E' - V(r)}{2mc^2} \right] (-i\hbar \underline{\underline{\sigma}} \cdot \nabla)^2 \underline{\Psi} - \frac{1}{2m} \frac{(-i\hbar \underline{\underline{\sigma}} \cdot \nabla V(r)) (-i\hbar \underline{\underline{\sigma}} \cdot \nabla) \underline{\Psi}}{2mc^2} + V(r) \underline{\Psi}$$

$$= \frac{\hbar^2}{2m} \left[1 - \frac{E' - V(r)}{2mc^2} \right] (\underline{\underline{\sigma}} \cdot \nabla)^2 \underline{\Psi} - \frac{\hbar^2}{4m^2 c^2} [\underline{\underline{\sigma}} \cdot \nabla V(r)] [\underline{\underline{\sigma}} \cdot \nabla] \underline{\Psi} + V(r) \underline{\Psi}$$

Now make use of few σ basic identities. \rightarrow

We make use of: $(\vec{\sigma} \cdot \nabla)^2 = \nabla^2$

$$(\vec{\sigma} \cdot \nabla V)(\vec{\sigma} \cdot \nabla \Psi) = (\nabla V) \cdot (\nabla \Psi) + i \vec{\sigma} \cdot [(\nabla V) \times (\nabla \Psi)]$$

Also, $\nabla V(r) = \frac{dV}{dr} \hat{r} \quad \therefore$ sph symm of $V(r)$

$$\therefore (\nabla V) \cdot (\nabla \Psi) = \frac{dV}{dr} \frac{\partial \Psi}{\partial r}$$

$$\begin{aligned} \therefore i \vec{\sigma} \cdot [(\nabla V) \times (\nabla \Psi)] &= i \vec{\sigma} \cdot \left(\frac{dV}{dr} \hat{r} \times \nabla \Psi \right) = -\vec{\sigma} \cdot \left(\frac{dV}{dr} \right) \frac{1}{r} \hat{r} \times \hat{p} \Psi \\ &= -\frac{1}{r} \left(\frac{dV}{dr} \right) \vec{\sigma} \cdot \hat{L} \Psi \end{aligned}$$

\therefore We have,

$$E' \Psi = \left[-\frac{\hbar^2}{2m} \nabla^2 + V(r) + \frac{\hbar}{2m} \frac{(E' - V(r))}{2mc^2} \nabla^2 - \frac{\hbar}{4m^2 c^2} \left\{ \frac{dV}{dr} \right\} \frac{\partial}{\partial r} - \frac{1}{r} \left(\frac{dV}{dr} \right) \vec{\sigma} \cdot \hat{L} \right] \Psi$$

$$E' \Psi = \left[\underbrace{-\frac{\hbar^2}{2m} \nabla^2 + V(r)}_{H_0} + \frac{\hbar}{4m^2 c^2} \frac{(E' - V(r))}{2m} \nabla^2 + \frac{\hbar}{4m^2 c^2} \frac{1}{r} \left(\frac{dV}{dr} \right) \vec{\sigma} \cdot \hat{L} - \frac{\hbar}{4m^2 c^2} \left(\frac{dV}{dr} \right) \frac{\partial}{\partial r} \right] \Psi$$

$$= \left[H_0 + \left(\frac{-1}{4m^2 c^2} \frac{\hbar^4}{2m} \right) + \frac{\hbar}{2m^2 c^2} \frac{1}{r} \left(\frac{dV}{dr} \right) \vec{S} \cdot \hat{L} - \frac{\hbar}{4m^2 c^2} \left(\frac{dV}{dr} \right) \frac{\partial}{\partial r} \right] \Psi$$

↓ KE correction
↓ Spin-orbit correction
↓ PE correction
↓ Darwin term

$$\vec{S} = \frac{\hbar \vec{\sigma}}{2} \Rightarrow (\vec{S} \cdot \vec{S}) \Psi = s(s+1) \hbar^2 \Psi$$

$s = \frac{1}{2}$

Spin angular momentum operator

$$\vec{S}^2 \Psi = \pm \frac{\hbar^2}{2} \Psi$$

$i=1, 3, 3$

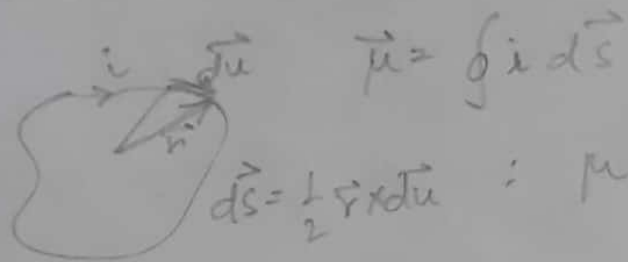
Recall Pauli $\vec{\sigma}_3$ and use \vec{S} :

$$E' \Psi = \left[\frac{1}{2m} (-i\hbar \nabla) \cdot (-i\hbar \nabla) + 2 \frac{e\hbar}{2m} \vec{S} \cdot \vec{B} + V(r) \right] \Psi$$

\therefore Intrinsic magnet moment: $\vec{\mu} = -\mu_B \vec{\sigma} = -g_s \frac{\mu_B}{\hbar} \vec{S} = -g_s \frac{e}{2m} \vec{S}$

g_s the gyromagnetic ratio

Angular momentum and magnetic moment



$$|\mu| = \frac{e}{2m} \hbar \sqrt{l(l+1)}$$

$$= \mu_B \sqrt{l(l+1)}$$

$$\hookrightarrow 9.27 \times 10^{-24} \text{ JT}^{-1}$$

$$\mu_z = -\frac{e \hbar m_l}{2m}$$

$$= -\mu_B m_l$$

Similarly for spin the Dirac theory gives

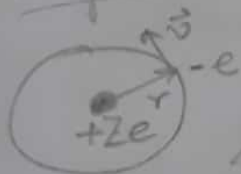
$$\mu_z = -g \mu_B m_s$$

$$g \downarrow$$

$$\rightarrow 2.0028192 \dots$$

for more accurate QED estimate when light is treated quantum mechanically

Magnetic field due to rotation motion of nucleus



Shift origin to centre



$$\vec{B} = \frac{\mu_0}{4\pi} \int \frac{d\vec{u} \times \vec{r}}{r^3}$$

Consider a circular loop with constant r

$$\therefore \int i d\vec{u} = \int \frac{dq}{dt} d\vec{u} = Ze \frac{d\vec{u}}{dt} = Ze(-\vec{v})$$

$$\therefore \text{for circular loop } \vec{B} = \frac{\mu_0}{4\pi} \frac{Ze}{r^3} (-\vec{v}) \times \vec{r} = \frac{\mu_0}{4\pi} \frac{Ze}{r^3} \vec{v} \times \vec{r}$$

$$\text{Now } \vec{E} = \frac{Ze}{4\pi\epsilon_0 r^2} \hat{r} = \frac{Ze}{4\pi\epsilon_0 r^3} \vec{r} \Rightarrow \vec{B} = \frac{\mu_0 Ze}{4\pi r^3} \times \frac{4\pi\epsilon_0 r^3}{Ze} \vec{E} \times \vec{v}$$

$$= \mu_0 \epsilon_0 \vec{E} \times \vec{v} = \frac{\vec{E} \times \vec{v}}{c^2}$$

\therefore Correction of energy of electron due to coupling of its intrinsic magnetic moment due to spin with the effective magnetic field due to rotation motion of nucleus around it: $-\mu_{spin} \cdot B_{field}$

$$= \left(-g_s \frac{\mu_B}{\hbar} \vec{S}\right) \cdot \left(\frac{\vec{E} \times \vec{v}}{c^2}\right) = \frac{1}{c^2} \left(-g_s \frac{\mu_B}{\hbar} \vec{S}\right) \cdot \left\{ \left(\frac{1}{c} \frac{d\vec{v}}{dt}\right) \times \vec{v} \right\}$$

$$= \frac{g_s \mu_B}{\hbar c^2} \left(\frac{1}{r} \frac{dV}{dr}\right) \vec{S} \cdot (\vec{v} \times \vec{r})$$

Almost \checkmark \uparrow \rightarrow can't be a book for only $v \ll c$ \rightarrow $10^8/3 \times 10^8$

$$= \frac{g_s}{2c^2 m^2} \left(\frac{1}{r} \frac{dv}{dr} \right) \vec{L} \cdot \vec{S}$$

Note that this is identical to what we got from Dirac theory except for the g_s which is 2. So a compensating 2 in the denominator must be there to be at par with Dirac theory and indeed such a factor exists. The imperfection in the above analysis is that the shift of origin is not an inertial translation but a non-inertial one since the rotating frame is non-inertial. As a result, an additional component of motion known as the Thomas Precession would arise which will take away half of the energy, leading to the desired 2 in the denominator.

Now let us revert back to our correction.

$$\text{Recall, } E(\psi) = \left[H_0 + \underbrace{\left(\frac{-1}{4m^2 c^2} \frac{\hbar^4}{2m} \right)}_{\text{KE correction}} + \underbrace{\frac{1}{2mc^2} \frac{1}{r} \left(\frac{dv}{dr} \right) \vec{L} \cdot \vec{S}}_{\text{L.S interaction}} - \underbrace{\frac{\hbar^2}{4m^2 c^2} \left(\frac{dv}{dr} \right) \frac{\partial^2}{\partial r^2}}_{\text{PE correction}} \right] \psi$$

There is a basic problem: four component ψ were normalized as $\int (\psi^\dagger \psi + \eta^\dagger \eta) dr = 1$

\therefore 2 component $\underline{\psi}$ is not normalized.

\therefore If with the correction term derived not not be Hermitian, indeed $\left[\frac{\partial}{\partial r} \right]$ is not Hermitian.

\therefore So the problem of normalization of $\underline{\psi}$ can be fixed by making the H Hermitian by adding the Hermitian counterpart of $\frac{\partial}{\partial r}$ to H. \rightarrow Proposed by Darwin.

So the PE correction is renamed as the Darwin term:

$$\frac{1}{2} \left[\left(-\frac{\hbar^2}{4m^2 c^2} \left(\frac{dv}{dr} \right) \frac{\partial^2}{\partial r^2} \right) + \left(-\frac{\hbar^2}{4m^2 c^2} \left(\frac{dv}{dr} \right) \frac{\partial}{\partial r} \right) \right]$$

$$\Rightarrow \int_{a_b} \psi^\dagger \psi = \left(-\frac{\hbar^2}{8m^2 c^2} \right) \left[\left[\left(\frac{dv}{dr} \frac{\partial}{\partial r} \right) \right]_{a_b} + \left[\left(\frac{dv}{dr} \frac{\partial}{\partial r} \right)^\dagger \right]_{a_b} \right]$$

$$\therefore D_{ab} = \left(-\frac{\hbar^2}{8m^2c^2} \right) \left[\langle a | \left(\frac{dv}{dr} \right) \frac{2}{2r} | b \rangle + \langle b | \left(\frac{dv}{dr} \right) \frac{2}{2r} | a \rangle \right]$$

$$\Rightarrow \left(-\frac{\hbar^2}{8m^2c^2} \right) \left[\int \psi_a^* \left(\frac{dv}{dr} \right) \nabla \psi_b \cdot d\mathbf{v} + \int \psi_b \left(\frac{dv}{dr} \right) \nabla \psi_a^* \cdot d\mathbf{v} \right]$$

$$= \left(-\frac{\hbar^2}{8m^2c^2} \right) \left[\dots + \dots + \int \psi_a^* \left(\nabla \left(\frac{dv}{dr} \right) \right) \psi_b \cdot d\mathbf{v} - \text{same} \right]$$

$$= \left(-\frac{\hbar^2}{8m^2c^2} \right) \left[\int \nabla \left[\psi_a^* \frac{dv}{dr} \psi_b \right] \cdot d\mathbf{v} - \int \psi_a^* \left(\nabla \left(\frac{dv}{dr} \right) \right) \psi_b \cdot d\mathbf{v} \right]$$

$$= \left(-\frac{\hbar^2}{8m^2c^2} \right) \left[\underbrace{\left[\psi_a^* \frac{dv}{dr} \psi_b \right]}_{\downarrow 0} - \int \psi_a^* \left(\nabla \left(\frac{dv}{dr} \right) \right) \psi_b \cdot d\mathbf{v} \right]$$

$$= -\frac{\hbar^2}{8m^2c^2} \left[\int \psi_a^* \left(\nabla \left(\frac{dv}{dr} \right) \right) \psi_b \cdot d\mathbf{v} \right]$$

with $V(r) = \frac{Ze^2}{4\pi\epsilon_0 r}$; $\left[\int \psi_a^* \left(\nabla \left(\frac{dv}{dr} \right) \right) \psi_b \cdot d\mathbf{v} \right] =$

$$= \left[\left(\frac{Ze^2}{4\pi\epsilon_0} \right) \int \psi_a^* \nabla \left(-\frac{1}{r} \right) \psi_b \cdot d\mathbf{v} \right]$$

$$= () \left[\int \nabla \left(\psi_a^* \frac{2}{2r} \psi_b \right) \cdot d\mathbf{v} - \int \frac{2}{2r} \left(\frac{1}{r} \right) \nabla \left(\psi_a^* \psi_b \right) \cdot r^2 \sin\theta \, d\theta \, d\phi \, dr \right]$$

$$= () \left[\dots + 4\pi \int_0^\infty \frac{2}{2r} \left(\psi_a^* \psi_b \right) dr \right] = 4\pi () \psi_a(r=0) \psi_b(r=0)$$

$$\therefore D_{ab} = \left(\frac{\hbar^2}{8m^2c^2} \right) \left(\frac{Ze^2}{4\pi\epsilon_0} \right) \langle a | \delta(r) | b \rangle$$

$$\therefore \text{P.E. contribution : Darwin term} = \frac{\hbar^2}{8m^2c^2} \left(\frac{Ze^2}{4\pi\epsilon_0} \right) \delta r$$

only contribute for $l=0$
 since there is a node
 at $r=0$ for $l > 0$

$$H = H_0 + H_1' + H_2' + H_3'$$

↓
↓
↘

Relativistic Contribution to KE Spin-orbit interaction Darwin term

$$\Psi_{nlm,ms} = \Psi_{nlm,ms} \begin{bmatrix} \chi_{\frac{1}{2},ms} \\ \chi_{\frac{1}{2},ms} \end{bmatrix}_{1 \times 2}$$

two component spinor

$$\chi_{\frac{1}{2},+\frac{1}{2}} = \begin{pmatrix} 1 \\ 0 \end{pmatrix}$$

$$\chi_{\frac{1}{2},-\frac{1}{2}} = \begin{pmatrix} 0 \\ 1 \end{pmatrix}$$

(spin orbitals or Pauli wave functions)

$$H_0 \Psi_{nlm,ms} = E_n \Psi_{nlm,ms} \rightarrow 2n^2 \text{ degenerate}$$

$m_s = \pm \frac{1}{2}$

We calculate 1st order correction due to H_1', H_2', H_3' with the spin orbitals as the zeroth order wavefn.

$$H_1' = -\frac{\hbar^4}{8m^3c^2} = -\frac{\hbar^4 \nabla^2 \nabla^2}{8m^3c^2} \rightarrow H_1' \text{ operates on } \Psi_{nlm,ms}$$

not the spinor part

It also turns out that H_1' commutes with L_z, S_z, L^2, S^2
 $\therefore H_1'$ is not going to mix states with specific l, m_l, m_s

Note that otherwise degenerate perturbation theory should have been required. \rightarrow Diagonal H_1' in degenerate basis

But H_1' is already diagonal in $\Psi_{nlm,ms}$

$$\Delta E = \langle \Psi_{nlm,ms} | -\frac{\hbar^4 \nabla^2 \nabla^2}{8m^3c^2} | \Psi_{nlm,ms} \rangle$$

$$= \frac{1}{2mc^2} \langle | \frac{\hbar^2}{2m} \frac{\hbar^2}{2m} | \rangle = \frac{1}{2mc^2} \langle | (H_0 + \frac{Ze^2}{4\pi\epsilon_0 r})^2 | \rangle$$

$$= \frac{1}{2mc^2} \left[\langle H_0 H_0 \rangle + \frac{Ze^2}{4\pi\epsilon_0} \langle \frac{1}{r} H_0 \rangle + \langle H_0 \frac{Ze^2}{4\pi\epsilon_0 r} \rangle + \left(\frac{Ze^2}{4\pi\epsilon_0} \right)^2 \langle \frac{1}{r^2} \rangle \right]$$

$$= \frac{1}{2mc^2} \left[E_n^2 + 2 \frac{Ze^2}{4\pi\epsilon_0} E_n \langle \frac{1}{r} \rangle_{nlm} + \left(\frac{Ze^2}{4\pi\epsilon_0} \right)^2 \langle \frac{1}{r^2} \rangle_{nlm} \right]$$

$$= -E_n \frac{(Ze)^2}{m^2} \left[\frac{3}{4} - \frac{n}{l + \frac{1}{2}} \right]$$

one interpretation of $\frac{1}{\alpha}$ is that it is the maximum charge Z which can support a stable cluster orbital around it: $\frac{mv^2}{r} = \frac{Ze^2}{4\pi\epsilon_0 r^2}$ $\rightarrow mv^2 = \frac{Ze^2}{4\pi\epsilon_0 r}$ $\rightarrow \alpha = \frac{e^2}{4\pi\epsilon_0 \hbar c}$

$\frac{1}{\alpha} = 137.0359991...$

$$\begin{aligned} \therefore \Delta E_2 &= \langle n l s j m_j | \hat{H}_2(r) | n l s j m_j \rangle \\ &= \int R_{nl}(r) [s j m_j]^\dagger \hat{H}_2(r) R_{nl}(r) [s j m_j] r^2 \sin \theta dr d\theta d\phi \end{aligned}$$

$$= \int |R_{nl}(r)|^2 \langle s j m_j | \hat{H}_2(r) | s j m_j \rangle r^2 dr$$

$$\langle s j m_j | \hat{H}_2(r) | s j m_j \rangle = \left(\frac{Ze^2}{4\pi\epsilon_0} \right) \frac{1}{2m_e c^2} \frac{Z^3}{a_0^3 n^3 l(l+\frac{1}{2})(l+1)} \text{ for } l \neq 0$$

divergent as $r \rightarrow \infty$ and $\psi(r \rightarrow \infty) \neq 0$ for $l=0$

$$\begin{aligned} \langle l s j m_j | \hat{L} \cdot \hat{S} | l s j m_j \rangle &= \frac{\hbar^2}{2} \langle l s j m_j | \hat{J}^2 - \hat{L}^2 - \hat{S}^2 | l s j m_j \rangle \\ j = l + \frac{1}{2} \Rightarrow \langle \hat{L} \cdot \hat{S} \rangle &= \frac{\hbar^2}{2} \left[(l + \frac{1}{2})(l + \frac{3}{2}) - l(l+1) - \frac{3}{4} \right] \\ &= \frac{\hbar^2}{2} l \Rightarrow 0 \text{ for } l=0 \\ j = l - \frac{1}{2} \Rightarrow \langle \hat{L} \cdot \hat{S} \rangle &= \frac{\hbar^2}{2} \left[(l - \frac{1}{2})(l + \frac{1}{2}) - l(l+1) - \frac{3}{4} \right] \\ &= \frac{\hbar^2}{2} (-l-1) \end{aligned}$$

$$\begin{aligned} \therefore \Delta E_2 &= \left(\frac{Ze^2}{4\pi\epsilon_0} \right) \frac{1}{2m_e c^2} \frac{Z^3 \hbar^2}{a_0^3 n^3} \left[\frac{1}{l(l+\frac{1}{2})(l+1)} \right] \times \begin{cases} l & \text{for } j=l+\frac{1}{2} \\ -l-1 & \text{for } j=l-\frac{1}{2} \end{cases} \\ &= \frac{1}{2} \left(\frac{e^2}{4\pi\epsilon_0} \right) \frac{1}{2m_e c^2} \frac{Z^4 \hbar^2}{\left(\frac{e^2}{4\pi\epsilon_0} \right)^3 \left(\frac{\hbar}{m} \right)^3 n^3} \left[\dots \right] \\ &= \frac{1}{4} \left(\frac{e^2}{4\pi\epsilon_0} \right)^4 \frac{Z^4}{\hbar^4 c^4} \frac{m_e c^2}{n^3} [\dots] \\ &= \frac{1}{4} Z^4 \left(\frac{e^2}{4\pi\epsilon_0 \hbar c} \right)^4 \frac{m_e c^2}{n^3} [\dots] \\ &= \frac{1}{4} Z^4 \alpha^4 \frac{m_e c^2}{n^3} [\dots] \\ &= -\frac{1}{2} \left[-\frac{1}{2} m_e c^2 \left(\frac{Z\alpha}{n} \right)^2 \right] Z^2 \alpha^2 \frac{1}{n} [\dots] \end{aligned}$$

Bohr radius a_0

$$\frac{m_e v}{a_0} = \frac{e^2}{4\pi\epsilon_0 a_0^2}$$

$$v = \frac{\hbar k}{m} = \frac{\hbar}{m} \frac{2\pi}{\lambda} = \frac{\hbar}{m} \frac{2\pi}{m \lambda}$$

$$= \frac{1}{m a_0}$$

$$\therefore \frac{m}{a_0} \left(\frac{\hbar}{m a_0} \right)^2 = \frac{e^2}{4\pi\epsilon_0 a_0^2}$$

$$\Rightarrow a_0 = \frac{4\pi\epsilon_0 \hbar^2}{m_e e^2}$$

Reduced $\alpha = \frac{e^2}{4\pi\epsilon_0 \hbar c}$

$$a_0^2 c^2 = \frac{e^2 \hbar^2}{4\pi\epsilon_0 m_e^2}$$

$$\Delta E_2 = -\frac{1}{2} E_n (Z\alpha)^2 \frac{1}{n} \frac{1}{l(l+\frac{1}{2})(l+1)} \left. \begin{array}{l} l \text{ for } j=l+\frac{1}{2} \\ -l-1 \text{ for } j=l-\frac{1}{2} \end{array} \right\} \times$$

E_n : Recall our interpretation of $\frac{1}{2}$ as the highest Z value which can sustain stable electron orbits with electron moving (almost) as fast as c

$$\therefore \frac{1}{2} \rightarrow c$$

$$\Rightarrow Z \rightarrow \alpha c Z$$

$$\therefore -\frac{1}{2} m v^2 = -\frac{1}{2} m c^2 (\alpha Z)^2 \text{ or more correctly as De Broglie quantization}$$

$$E_n = -\frac{1}{2} m c^2 \left(\frac{\alpha Z}{n}\right)^2$$

Remember that $\langle \frac{1}{r} \rangle \propto \frac{1}{a_0}$ for $l=0$ (although it appears from above that it may contribute for $l>0$ as well)

$$\Rightarrow \Delta E_2 = 0 \text{ for } l=0$$

Darwin term: $\Delta E_3 = \langle \psi_{n00} | \frac{\hbar^2}{2m^2 c^2} \frac{\partial^2}{\partial x^2} \psi_{n00} \rangle$

(already argued that it is valid for $l=0$ only)

$$= \frac{\hbar^2}{2m^2 c^2} \frac{Z e^2}{4\pi \epsilon_0} |\psi_{n00}(r=0)|^2$$

$$= \frac{\hbar^2}{2m^2 c^2} \frac{Z e^2}{4\pi \epsilon_0} \frac{4 Z^3}{n^3 a_0^3} \frac{1}{4\pi}$$

$$= \left(\frac{e^2}{4\pi \epsilon_0}\right) \frac{\hbar^2}{2m^2 c^2} \frac{Z^4}{\left(\frac{e^2}{4\pi \epsilon_0}\right)^{-3} \left(\frac{\hbar^2}{m}\right)^3} \frac{1}{n^3}$$

$$= \left(\frac{e^2}{4\pi \epsilon_0}\right)^4 \frac{Z^4}{\hbar^4 c^2} \frac{1}{2} \frac{1}{n^3}$$

$$= (\alpha Z)^4 \frac{m c^2}{2} \frac{1}{n^3}$$

$$= \frac{1}{2} m c^2 \left(\frac{Z\alpha}{n}\right)^2 (\alpha Z)^2 \frac{1}{n}$$

$$= -E_n (\alpha Z)^2 \frac{1}{n} \text{ only for } l=0$$

Recall $R_{10} = 2 \left(\frac{Z}{a_0}\right)^{3/2} e^{-Zr/a_0}$

$R_{20} = 2 \left(\frac{Z}{a_0}\right)^{3/2} \left(1 - \frac{Zr}{2a_0}\right) e^{-Zr/2a_0}$

$\frac{1}{\sqrt{4\pi}}$ will come for ψ_{00}

Net correction $\Delta E_{\text{HT}} = \Delta E_1 + \Delta E_2 + \Delta E_3$

$$\Delta E_{\text{HT}} = -\frac{E_n (\alpha Z)^2}{n^2} \left[\left(\frac{3}{4} - \frac{n}{2}\right) + 0 + n \right] = -\frac{E_n (\alpha Z)^2}{n^2} \left[\frac{3}{4} - n \right]$$

$$= \frac{E_n (\alpha Z)^2}{n^2} \left[\frac{n}{j+\frac{1}{2}} - \frac{3}{4} \right]$$

for $l \geq 1, j = l + \frac{1}{2}$

$$\begin{aligned} \Delta E_{for} &= \Delta E_1 + \Delta E_2 + \Delta E_3 \quad \left\{ \begin{array}{l} \text{since Darwin term} \\ \text{contributes only for } l=0 \end{array} \right. \\ &= -\frac{E_n (\alpha Z)^2}{n^2} \left[\frac{3}{4} - \frac{n}{l + \frac{1}{2}} + \frac{n l}{2l(l+1)(l + \frac{1}{2})} \right] \\ &= -\frac{E_n (\alpha Z)^2}{n^2} \left[\frac{3}{4} - \frac{2nl(l+1) - nl}{2l(l+1)(l + \frac{1}{2})} \right] \\ &= -\frac{E_n (\alpha Z)^2}{n^2} \left[\frac{3}{4} - \frac{2nl(l + \frac{1}{2})}{2l(l+1)(l + \frac{1}{2})} \right] \\ &= \frac{E_n (\alpha Z)^2}{n^2} \left[\frac{n}{j + \frac{1}{2}} - \frac{3}{4} \right] \end{aligned}$$

for $l \geq 1, j = l - \frac{1}{2}$

$$\begin{aligned} \Delta E_{for} &= -\frac{E_n (\alpha Z)^2}{n^2} \left[\frac{3}{4} - \frac{n}{l + \frac{1}{2}} - \frac{n(l+1)}{2l(l+1)(l + \frac{1}{2})} \right] \\ &= -\frac{E_n (\alpha Z)^2}{n^2} \left[\frac{3}{4} - \frac{2nl(l+1) + n(l+1)}{2l(l+1)(l + \frac{1}{2})} \right] \\ &= -\frac{E_n (\alpha Z)^2}{n^2} \left[\frac{3}{4} - \frac{2n(l+1)(l + \frac{1}{2})}{2l(l+1)(l + \frac{1}{2})} \right] \\ &= \frac{-E_n (\alpha Z)^2}{n^2} \left[\frac{3}{4} - \frac{n}{l} \right] \\ &= \frac{E_n (\alpha Z)^2}{n^2} \left[\frac{n}{j + \frac{1}{2}} - \frac{3}{4} \right] \end{aligned}$$

$\therefore \Delta E_{for} = \frac{E_n (\alpha Z)^2}{n^2} \left[\frac{n}{j + \frac{1}{2}} - \frac{3}{4} \right]$ for all values of l .

$\therefore E_{nl} \xrightarrow{\text{after fine structure correction}} E_{nj} = E_n \left[1 + \frac{(\alpha Z)^2}{n^2} \left(\frac{n}{j + \frac{1}{2}} - \frac{3}{4} \right) \right]$

-ve $\quad E_{nj} < E_{nl}$ as \nearrow

Note for $n=1, l=0, j = \frac{1}{2} \Rightarrow \frac{n}{j + \frac{1}{2}} = 1$ binding of electron increases.

$n=2 \Rightarrow l_{max} = 1 \Rightarrow j_{max} = \frac{3}{2} \Rightarrow \left(\frac{n}{j + \frac{1}{2}} \right)_{min} = 1; \left(\frac{n}{j + \frac{1}{2}} \right)_{max} = 2$ $L = \hbar \sqrt{l(l+1)}$